



# Behavior in time of solutions to a degenerate chemotaxis system with flux limitation

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## ABSTRACT

We study a new class of Keller–Segel models, which presents a limited flux and an optimal transport of cells density according to chemical signal density. As a prototype of this class we study radially symmetric solutions to the parabolic–elliptic system

$$\begin{cases} u_t = \nabla \cdot \left( \frac{u \nabla u}{\sqrt{u^2 + |\nabla u|^2}} \right) - \chi k_f \nabla \cdot \left( \frac{u \nabla v}{(1 + |\nabla v|^2)^{\alpha}} \right), & x \in \Omega, t > 0, \\ 0 = \Delta v - \mu + u, & x \in \Omega, t > 0 \end{cases}$$

under no flux boundary conditions in a ball  $B = \Omega \subset \mathbb{R}^N$  and initial condition  $u(x, 0) = u_0(x) > 0$ ,  $\chi > 0$ ,  $\alpha > 0$ ,  $k_f > 0$  and  $\mu = \frac{1}{|\Omega|} \int_{\Omega} u_0 dx$ . Under suitable conditions on  $\alpha$  and  $u_0$  it is shown that the solution blows up in  $L^\infty$ -norm at a finite time  $T_{max}$  and for some  $p > 1$  it blows up also in  $L^p$ -norm. The proofs are mainly based on an helpful change of variables, on comparison arguments and some suitable estimates.

## 1. Introduction

Let us consider the chemotaxis system with nonlinear diffusion and flux limitation,

$$\begin{cases} u_t = \nabla \cdot \left( \frac{u \nabla u}{\sqrt{u^2 + |\nabla u|^2}} \right) - \chi \nabla \cdot (u f(|\nabla v|^2) \nabla v), & x \in \Omega, t > 0, \\ 0 = \Delta v - \mu + u, & x \in \Omega, t > 0, \\ \frac{u \nabla u \cdot \nu}{\sqrt{u^2 + |\nabla u|^2}} = \nabla v \cdot \nu = 0, & x \in \partial \Omega, t > 0, \\ u(x, 0) = u_0(x) > 0, & x \in \Omega, \end{cases} \quad (1.1)$$

with  $\Omega$  a ball in  $\mathbb{R}^N$ ,  $N \geq 3$ , the constant  $\chi > 0$ ,  $\int_{\Omega} v dx = 0$ , the initial data  $u_0$  such that

$$u_0 \in C^2(\overline{\Omega}), \text{ radially symmetric and positive in } \overline{\Omega}, \quad (1.2)$$

where

$$\mu := \frac{1}{|\Omega|} \int_{\Omega} u_0(x) dx \quad (1.3)$$

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and

$$f(|\nabla v|^2) = k_f(1 + |\nabla v|^2)^{-\alpha} \tag{1.4}$$

with some  $k_f > 0$  and  $\alpha > 0$ .

We point out that the general structure of the model is

$$\begin{cases} u_t = \nabla \cdot (D_u(u, v)\nabla u) - \nabla \cdot (S(u, v)u\nabla v) + H_1(u, v), \\ \tau v_t = D_v \Delta v + H_2(u, v), \end{cases} \tag{1.5}$$

where  $u(x, t)$  represents the cells density,  $v(x, t)$  is the density of the chemoattractant,  $S$  measures the chemotactic sensitivity,  $D_u, D_v$  are two positive functions, representing the diffusivity of the cells and of the chemoattractant respectively,  $H_1, H_2$  model source terms related to interactions (see [1]).

The most investigated model has  $D_u = S = 1, H_1 = 0$  or  $H_1$  of logistic type and  $H_2 = -v + u$  (see [2] and the survey paper [3]). Interesting results have been established also for the parabolic–elliptic chemotaxis systems ( $\tau = 0$ ).

More recently nonlinear diffusion terms have been considered with  $D$  and  $S$  depending not only on  $u$  and  $v$ , but also on their gradient.

• Winkler and Djie in [4], with the aim to study simplified models in the theoretical description of chemotaxis phenomena under the influence of the volume-filling effect, considered (1.5) with  $D_u = (u + 1)^{-p}, S = (u + 1)^{q-1}, D_v = 1, H_2 = u - \mu$ , and proved that if  $p + q < 2N$  then all solutions are global in time and bounded, whereas if  $p + q > 2N, q > 0$ , and  $\Omega$  is a ball, then there exist solutions that become unbounded in finite time.

• In [5] (1.5) is investigated in  $\Omega \times (0, T)$  bounded in  $\mathbb{R}^N$  with  $D_u = (u + \alpha)^{m_1-1}$  and  $S = \chi(u + \alpha)^{m_2-2}, D_v = 1, H_2 = u - \mu$ , under Neumann boundary conditions and initial conditions,  $\alpha > 0, \chi > 0, m_1, m_2 \in \mathbb{R}$ . It is proved that for some  $p_0 > \frac{N}{2}(m_2 - m_1)$  any blowing up solution in  $L^\infty(\Omega)$ -norm, blows up also in  $L^{p_0}(\Omega)$ -norm and the blow-up time is estimated.

Recently the case  $S = S(|\nabla v|)$  depending on the gradient of  $v$  (flux limitation term) received considerable attention in the biomathematical literature.

Here we report only the most important results on flux limitation.

- If  $D_u = D_v = 1$  and  $S = \chi|\nabla v|^{p-2}, \chi > 0, H_2 = u - \mu, \Omega \subset \mathbb{R}^N,$

$$p \in (1, \infty) \text{ if } N = 1; \quad p \in \left(1, \frac{N}{N-1}\right) \text{ if } N \geq 2,$$

Negreanu and Tello in [6] obtained uniform bounds in  $L^\infty(\Omega)$  and existence of global in time solutions of (1.5) with  $\tau = 0$ ; moreover they prove that for the one-dimensional case there exist infinitely many non-constant steady-states for  $p \in (1, 2)$ . As to its parabolic–parabolic case ( $\tau = 1$ ), Yan and Li in [7] obtained global existence of weak solutions which are uniformly bounded provided that  $1 < p < \frac{N}{N-1}$ .

In [8], Kohatsu and Yokota established stability of constant equilibria for small initial data and  $p \in (1, \infty)$ .

- If  $D_u = \frac{u}{\sqrt{u^2 + |\nabla u|^2}}$  and  $S = \frac{1}{\sqrt{1 + |\nabla v|^2}}, H_1 = 0, H_2 = u - \mu,$  Bellomo and Winkler [1] obtained global existence of bounded

classical solutions for arbitrary positive radial initial data  $u_0 \in C^3(\bar{\Omega})$  when either

$$N \geq 2 \text{ and } \chi < 1, \text{ or } N = 1, \chi > 0 \text{ and } m := \int_{\Omega} u_0 < m_c,$$

$$\text{with } m_c := \frac{1}{\sqrt{\chi^2 - 1}}, \text{ if } \chi > 1, \text{ and } m_c := \infty \text{ if } \chi \leq 1.$$

In [9], the authors showed that the above conditions are essentially optimal in the sense that if  $\chi > 1$  and

$$m > \frac{1}{\sqrt{\chi^2 - 1}}, \text{ if } N = 1; \quad m > 0 \text{ arbitrary, if } N \geq 2,$$

there exists  $u_0 \in C^3(\bar{\Omega})$  with  $\int_{\Omega} u_0 = m$ , such that for some  $T > 0$  there exists a unique classical solution blowing up at time  $T$  in  $L^\infty(\Omega)$ -norm.

• If  $D_u = 1$  and  $S(|\nabla v|^2) \geq K_f(1 + |\nabla v|^2)^{-\alpha}, K_f > 0, \chi = 1, 0 < \alpha < \frac{N-2}{2(N-1)}, \Omega$  a ball in  $\mathbb{R}^N$ , with  $N \geq 3$ , for a considerably large set of radially symmetric initial data, Winkler in [10] proved that the problem admits solutions blowing up in finite time in  $L^\infty$ -norm for the first component. Otherwise, if  $S(|\nabla v|^2) \leq K_f(1 + |\nabla v|^2)^{-\alpha}, \chi = 1$  and  $\alpha$  satisfies

$$\begin{cases} \alpha > \frac{N-2}{2(N-1)}, \text{ for } N \geq 2, \\ \alpha \in \mathbb{R}, \text{ for } N = 1, \end{cases}$$

in general (not symmetric setting), no blow-up solution can exist then a global bounded solution exists for arbitrary nonnegative continuous initial data.

The case  $\alpha = \frac{N-2}{2(N-1)}$  plays the role of a critical exponent and it is still an open problem.

• If  $D_u = 1$  and  $S(|\nabla v|^2) = K_f(1 + |\nabla v|^2)^{-\alpha}, K_f > 0, \chi = 1, 0 < \alpha < \frac{N-2}{2(N-1)}, \Omega = B_R(0) \subset \mathbb{R}^N$ , with  $N \geq 3$ , Marras et al. in [11], for suitable initial data, proved that a solution which blows up in  $L^\infty$ -norm blows up also in  $L^p$ -norm for some  $p > \frac{N}{2}$ . Moreover, a safe time interval of existence of the solution  $[0, T]$  is obtained, with  $T$  a lower bound of the blow-up time. Moreover the same authors in [12] extended the results in [11] when a source term of logistic type is present in the first equation.

• If  $D_u = \frac{u^p}{\sqrt{u^2 + |\nabla u|^2}}$  and  $S(u, |\nabla v|^2) = \chi \frac{u^{q-1}}{\sqrt{1 + |\nabla v|^2}}$ , Chiyoda et al. in [13] considered the system (1.5) in a ball in  $\mathbb{R}^N$ ,  $N \in \mathbb{N}$ , under no-flux boundary conditions and initial condition  $u_0(x) > 0$ . Assuming suitable conditions for  $\chi$  and  $u_0$  when  $1 \leq p \leq q$ , they obtained existence of blow-up solutions. When  $p = q = 1$ , the system reduces to (1.1) when  $\alpha = \frac{1}{2}$ .

If in (1.5), the source  $H_1 \neq 0$ , many interesting results on local, global existence and blow up of solutions have been derived (see for instance [14–17] and references therein).

**Main Results** The present work is addressed to study the behavior in time of the solutions of (1.1) where the diffusion term is nonlinear, depending on  $u$  and  $\nabla u$ , as well as the drift term, depending on  $v$  and  $\nabla v$ .

In particular in Section 2 we construct an initial data such that the solution of problem (1.1) blows up in  $L^\infty$ -norm. We prove that the solutions of (1.1) blow up at finite time in  $L^p$ -norm, for some  $p > 1$ , if they blow up in  $L^\infty$ -norm (Section 3).

For the proof of our results we use suitable change of variables, comparison arguments and some helpful estimates.

We want to prove the following theorems.

**Theorem 1.1 (Finite-time blow-up in  $L^\infty$ -norm).** Let  $\Omega = B_R(0) \subset \mathbb{R}^N$ ,  $N \geq 3$  and  $R > 0$ ,  $f$  satisfy (1.4). Assume

$$0 < \alpha < \frac{N - 2}{2(N - 1)}. \tag{1.6}$$

Then for all  $\mu > 0$  there exists a radially symmetric positive initial data  $u_0 \in C^\infty(\bar{\Omega})$  such that satisfies (1.3), and such that there exists a radially symmetric positive classical solution  $(u, v)$  of (1.1) in  $\Omega \times (0, T_{max})$  for some  $T_{max} > 0$  which blows up at  $T_{max}$  in the sense that  $\limsup_{t \nearrow T_{max}} \|u(\cdot, t)\|_{L^\infty(\Omega)} = \infty$ .

**Theorem 1.2 (Finite-time blow-up in  $L^p$ -norm).** Let  $\Omega = B_R(0) \subset \mathbb{R}^N$ ,  $N \geq 3$  and  $R > 0$ . Then, a classical solution  $(u, v)$  of (1.1), provided by Theorem 1.1, is such that for all  $p > N$ ,

$$\limsup_{t \nearrow T_{max}} \|u(\cdot, t)\|_{L^p(\Omega)} = \infty.$$

## 2. Blow-up in $L^\infty$ -norm

Before discussing blow-up in  $L^\infty$ -norm, we give a result on local existence of classical solutions to (1.1). We note that blow-up criterion via the standard manner (see [18, Lemma 2.1]) says that

$$\text{if } T_{max} < \infty, \text{ then either } \liminf_{t \nearrow T_{max}} \inf_{x \in \Omega} u(x, t) = 0 \text{ or } \limsup_{t \nearrow T_{max}} \|u(\cdot, t)\|_{W^{1,\infty}(\Omega)} = \infty.$$

This includes possibility of extinction and gradient blow-up of solutions, whereas the following lemma presents a simple criterion ruling out this possibility.

**Lemma 2.1.** Let  $\Omega = B_R(0) \subset \mathbb{R}^N$ ,  $N \geq 1$  and  $R > 0$ ,  $f$  satisfy (1.4) with  $\alpha > 0$ . Assume  $u_0$  complies (1.2) and (1.3). Then there exist  $T_{max} \in (0, \infty]$  and a pair  $(u, v)$  of positive radially symmetric functions  $u \in C^{2,1}(\bar{\Omega} \times [0, T_{max}))$  and  $v \in C^{2,0}(\bar{\Omega} \times [0, T_{max}))$  which solve (1.1) classically in  $\Omega \times (0, T_{max})$ , and moreover  $u$  satisfies that

$$\text{if } T_{max} < \infty, \text{ then } \limsup_{t \nearrow T_{max}} \|u(\cdot, t)\|_{L^\infty(\Omega)} = \infty.$$

**Proof.** The claim can be proved by a pointwise lower estimate for  $u$  and by a uniform estimate for  $|\nabla u|$  almost in the same way as in [18].  $\square$

In order to obtain the blow-up phenomena in finite time of  $L^\infty$ -norm of solutions on (1.1), we follow the proof of Theorem 1.1 in [11].

So we first transform the system (1.1) in a non local scalar parabolic equation.

**Transformation of (1.1) in nonlocal scalar parabolic equation.**

Assume  $\Omega = B_R(0) \subset \mathbb{R}^N$ ,  $N \geq 3$ ,  $R > 0$  and  $u_0 \in C^0(\bar{\Omega})$  is radially symmetric with respect to  $x = 0$ . If  $(u, v)$  is the corresponding radial solution in  $\Omega \times (0, T_{max})$ , we write  $u = u(r, t)$  and  $v = v(r, t)$  with  $r = |x| \in [0, R]$ .

Following Jäger–Luckhaus [19] we introduce the mass accumulation function

$$w(s, t) := \int_0^s \frac{1}{r^{N-1}} \rho^{N-1} u(\rho, t) d\rho, \quad s = r^N \in [0, R^N], \quad t \in [0, T_{max}). \tag{2.1}$$

Then we have

$$w_s(s, t) = \frac{1}{N} u(s^{\frac{1}{N}}, t) > 0, \quad w_{ss}(s, t) = \frac{1}{N^2} s^{\frac{1}{N}-1} u_r(s^{\frac{1}{N}}, t).$$

From the second equation in (1.1) we deduce

$$\frac{1}{r^{N-1}} (r^{N-1} v_r(r, t))_r = \mu - u(r, t), \quad (r^{N-1} v_r(r, t))_r = \mu r^{N-1} - u(r, t) r^{N-1} \tag{2.2}$$

and by integrating (2.2),

$$r^{N-1}v_r(r, t) = \mu \int_0^r \rho^{N-1}d\rho - \int_0^r \rho^{N-1}u(\rho, t)d\rho = \frac{\mu r^N}{N} - w.$$

Using (1.1), we obtain

$$\begin{aligned} w_t(s, t) &= \int_0^{s^{\frac{1}{N}}} r^{N-1}u_t(r, t)dr \\ &= \int_0^{s^{\frac{1}{N}}} \left( r^{N-1} \frac{uu_r}{\sqrt{u^2 + u_r^2}} \right)_r dr - \chi \int_0^{s^{\frac{1}{N}}} \left( r^{N-1}u(r, t)v_r f(v_r^2) \right)_r dr \\ &= s^{1-\frac{1}{N}} \frac{uu_r}{\sqrt{u^2 + u_r^2}} - \chi k_f s^{1-\frac{1}{N}} \frac{uv_r}{(1 + v_r^2)^\alpha}. \end{aligned}$$

Since  $u = N w_s$  and  $u_r = N^2 s^{1-\frac{1}{N}} w_{ss}$ , we derive

$$\begin{aligned} w_t(s, t) &= s^{1-\frac{1}{N}} \frac{Nw_s N^2 s^{1-\frac{2}{N}} w_{ss}}{\sqrt{N^2 w_s^2 + N^4 s^{2-\frac{2}{N}} w_{ss}^2}} - \chi k_f \frac{Nw_s (\frac{\mu}{N}s - w)}{(1 + (\frac{\mu}{N}s^{\frac{1}{N}} - s^{\frac{1}{N}-1}w)^2)^\alpha} \\ &= N^2 \frac{s^{2-\frac{2}{N}} w_s w_{ss}}{\sqrt{w_s^2 + N^2 s^{2-\frac{2}{N}} w_{ss}^2}} + \chi k_f \cdot N \frac{w_s (w - \frac{\mu}{N}s)}{(1 + (s^{\frac{1}{N}-1}w - \frac{\mu}{N}s^{\frac{1}{N}})^2)^\alpha} \end{aligned}$$

for  $s \in (0, R^N)$  and  $t \in (0, T_{max})$ . Thus we can conclude that  $w$  satisfies the parabolic problem

$$\begin{cases} w_t = N^2 \frac{s^{2-\frac{2}{N}} w_s w_{ss}}{\sqrt{w_s^2 + N^2 s^{2-\frac{2}{N}} w_{ss}^2}} + \chi k_f \cdot N \frac{w_s (w - \frac{\mu}{N}s)}{(1 + (s^{\frac{1}{N}-1}w - \frac{\mu}{N}s^{\frac{1}{N}})^2)^\alpha}, \\ w(0, t) = 0, \quad w(R^N, t) = \frac{\mu R^N}{N}, \quad t \in (0, T_{max}), \\ w(s, 0) = w_0(s), \quad s \in (0, R^N) \end{cases} \tag{2.3}$$

with

$$w_0(s) = \int_0^{s^{\frac{1}{N}}} \rho^{N-1}u_0(\rho)d\rho, \quad s \in [0, R^N].$$

The next step is to prove that the functional  $\int_0^{R^N} s^{-a}w^b(s, t)ds$ , with  $b \in (0, 1)$ , satisfies a differential inequality.

To this end we first prove the following lemma.

**Lemma 2.2.** Assume  $\Omega = B_R(0) \subset \mathbb{R}^N$  with some  $R > 0$  and  $N \geq 2$ . Let  $u_0 \in C^0(\overline{\Omega})$  be radial and let  $(u, v)$  be the solution of (1.1). Then for all  $a > 0$ ,  $b \in (0, 1)$  and  $\epsilon > 0$  the function  $w$  defined in (2.1) satisfies

$$\begin{aligned} &\frac{1}{b} \frac{d}{dt} \int_0^{R^N} (s + \epsilon)^{-a} w^b(s, t) ds \\ &\geq -N \int_0^{R^N} s^{1-\frac{1}{N}} (s + \epsilon)^{-a} w^{b-1} w_s ds \\ &\quad + N k_f \bar{C} \int_0^{R^N} s^{(2-\frac{2}{N})\alpha} (s + \epsilon)^{-a} w^b w_s ds - \mu \chi k_f \int_0^{R^N} s (s + \epsilon)^{-a} w^{b-1} w_s ds \end{aligned} \tag{2.4}$$

for all  $t \in (0, T_{max})$  and  $\bar{C} := \bar{C}(R, N, \mu, \alpha)$  a positive constant.

**Proof.** Following Lemma 3.2 in [12], we multiply the first equation in (2.3) by  $(s + \epsilon)^{-a}w^{b-1}(s, \tau)$  and integrating from 0 to  $R^N$  we obtain

$$\begin{aligned} &\frac{1}{b} \frac{d}{dt} \int_0^{R^N} (s + \epsilon)^{-a} w^b(s, t) ds \\ &= N^2 \int_0^{R^N} s^{2-\frac{2}{N}} (s + \epsilon)^{-a} w^{b-1} \frac{w_s w_{ss}}{\sqrt{w_s^2 + N^2 s^{2-\frac{2}{N}} w_{ss}^2}} ds \\ &\quad + N \chi k_f \int_0^{R^N} (s + \epsilon)^{-a} w^{b-1} \frac{w_s (w - \frac{\mu}{N}s)}{[1 + s^{\frac{2}{N}-2}(w - \frac{\mu}{N}s)^2]^\alpha} ds \end{aligned}$$

$$=: I_1 + I_2. \tag{2.5}$$

To estimate  $I_1$ , since  $w_s \geq 0$ , we observe that  $w_s w_{ss} \geq -|w_{ss}|w_s$  and thus we obtain

$$\frac{w_s w_{ss}}{\sqrt{w_s^2 + N^2 s^{2-\frac{2}{N}} w_{ss}^2}} \geq \frac{-w_s |w_{ss}|}{N s^{1-\frac{1}{N}} |w_{ss}|}.$$

Then we can write

$$\begin{aligned} I_1 &= N^2 \int_0^{R^N} s^{2-\frac{2}{N}} (s + \epsilon)^{-a} w^{b-1} \frac{w_s w_{ss}}{\sqrt{w_s^2 + N^2 s^{2-\frac{2}{N}} w_{ss}^2}} ds \\ &\geq -N \int_0^{R^N} s^{1-\frac{1}{N}} (s + \epsilon)^{-a} w^{b-1} w_s ds. \end{aligned} \tag{2.6}$$

We now consider  $I_2$  and we obtain that

$$\begin{aligned} I_2 &= N \chi k_f \int_0^{R^N} (s + \epsilon)^{-a} w^{b-1} \frac{w_s (w - \frac{\mu}{N} s)}{[1 + s^{\frac{2}{N}-2} (w - \frac{\mu}{N} s)^2]^\alpha} ds \\ &= N \chi k_f \int_0^{R^N} (s + \epsilon)^{-a} w^b \frac{w_s}{[1 + s^{\frac{2}{N}-2} (w - \frac{\mu}{N} s)^2]^\alpha} ds \\ &\quad - \mu \chi k_f \int_0^{R^N} s (s + \epsilon)^{-a} w^{b-1} \frac{w_s}{[1 + s^{\frac{2}{N}-2} (w - \frac{\mu}{N} s)^2]^\alpha} ds \\ &=: I_{21} + I_{22}. \end{aligned} \tag{2.7}$$

To estimate the term  $I_{21}$  we observe that since  $w \leq \frac{\mu}{N} R^N$  and  $s \leq R^N$  we obtain  $(w - \frac{\mu}{N} s)^2 \leq w^2 + \frac{\mu^2}{N^2} s^2 \leq 2 \frac{\mu^2}{N^2} R^{2N}$ . So we can write

$$\frac{1}{[1 + s^{\frac{2}{N}-2} (w - \frac{\mu}{N} s)^2]^\alpha} = \frac{s^{(2-\frac{2}{N})\alpha}}{[s^{2-\frac{2}{N}} + (w - \frac{\mu}{N} s)^2]^\alpha} \geq \frac{s^{(2-\frac{2}{N})\alpha}}{[R^{2N-2} + 2 \frac{\mu^2}{N^2} R^{2N}]^\alpha} =: \bar{C} s^{(2-\frac{2}{N})\alpha},$$

with  $\bar{C} = \frac{1}{[R^{2N-2} + 2 \frac{\mu^2}{N^2} R^{2N}]^\alpha}$ . By substituting the previous estimate in  $I_{21}$  we arrive at

$$I_{21} \geq N k_f \bar{C} \int_0^{R^N} s^{(2-\frac{2}{N})\alpha} (s + \epsilon)^{-a} w^b w_s ds. \tag{2.8}$$

Now we estimate  $I_{22}$ .

Since  $\frac{1}{[1 + s^{\frac{2}{N}-2} (w - \frac{\mu}{N} s)^2]^\alpha} \leq 1$ , we obtain

$$\begin{aligned} I_{22} &= -\mu \chi k_f \int_0^{R^N} s (s + \epsilon)^{-a} w^{b-1} \frac{w_s}{[1 + s^{\frac{2}{N}-2} (w - \frac{\mu}{N} s)^2]^\alpha} ds \\ &\geq -\mu \chi k_f \int_0^{R^N} s (s + \epsilon)^{-a} w^{b-1} w_s ds. \end{aligned} \tag{2.9}$$

Substituting (2.6), (2.7) (with (2.8) and (2.9)) in (2.5) we obtain (2.4).  $\square$

In order to obtain a differential inequality for the functional  $\int_0^{R^N} s^{-a} w^b(s, t) dt$ , we take into account of Lemma 2.2 and we prove the following lemma.

**Lemma 2.3.** *Suppose that  $\alpha$  fulfills (1.6). Then there exist  $a > 0$ ,  $b \in (0, 1)$  satisfying*

$$\begin{aligned} 1 &< a < 2(1 - \alpha) \frac{N - 1}{N}, \\ b_0 &< b < 1, \quad b_0 := \frac{\frac{N}{N-1} a - 1}{1 - 2\alpha}. \end{aligned} \tag{2.10}$$

and some positive  $\delta_1(a, b, \mu, \chi, k_f)$ ,  $\delta_2(a, b, N, k_f)$  and  $\delta_3(a, b, R, N, \mu, \chi, k_f)$  such that

$$\begin{aligned} \frac{1}{b} \int_0^{R^N} s^{-a} w^b(s, t) ds &\geq \frac{1}{b} \int_0^{R^N} s^{-a} w_0^b(s) ds - \delta_1 \int_0^t \int_0^{R^N} s^{-a} w^b(s, \tau) ds d\tau \\ &\quad + \delta_2 \int_0^t \left( \int_0^{R^N} s^{-a} w^b(s, \tau) ds \right)^{\frac{b+1}{b}} d\tau - \delta_3 t. \end{aligned} \tag{2.11}$$

**Proof.** We recall the inequality (2.4):

$$\begin{aligned} \frac{1}{b} \frac{d}{dt} \int_0^{R^N} (s + \epsilon)^{-a} w^b(s, t) ds &\geq -N \int_0^{R^N} s^{1-\frac{1}{N}} (s + \epsilon)^{-a} w^{b-1} w_s ds \\ &\quad + N k_f \bar{C} \int_0^{R^N} s^{2(1-\frac{1}{N})\alpha} (s + \epsilon)^{-a} w^b w_s ds \\ &\quad - \mu \chi k_f \int_0^{R^N} s (s + \epsilon)^{-a} w^{b-1} w_s ds \\ &=: I_1 + I_2 + I_3. \end{aligned} \tag{2.12}$$

Integration by parts of  $I_1$  of (2.12) leads to

$$\begin{aligned} I_1 &= -N \int_0^{R^N} s^{1-\frac{1}{N}} (s + \epsilon)^{-a} w^{b-1} w_s ds \\ &= -\frac{N}{b} [s^{1-\frac{1}{N}} (s + \epsilon)^{-a} w^b]_0^{R^N} + \frac{N}{b} \int_0^{R^N} \frac{d}{ds} (s^{1-\frac{1}{N}} (s + \epsilon)^{-a}) w^b ds \\ &= -\frac{N}{b} R^{N-1} (R^N + \epsilon)^{-a} w^b(R^N, t) + \frac{N}{b} \int_0^{R^N} \frac{d}{ds} (s^{1-\frac{1}{N}} (s + \epsilon)^{-a}) w^b ds \\ &= -\frac{N}{b} R^{N-1} (R^N + \epsilon)^{-a} \frac{\mu^b R^{bN}}{N^b} + \frac{N}{b} \int_0^{R^N} \frac{d}{ds} (s^{1-\frac{1}{N}} (s + \epsilon)^{-a}) w^b ds \\ &= -A(\epsilon) + \frac{N}{b} (1 - \frac{1}{N}) \int_0^{R^N} s^{-\frac{1}{N}} (s + \epsilon)^{-a} w^b ds - a \frac{N}{b} \int_0^{R^N} s^{1-\frac{1}{N}} (s + \epsilon)^{-a-1} w^b ds, \end{aligned}$$

with  $A(\epsilon) = \frac{N}{b} R^{N-1} (R^N + \epsilon)^{-a} \frac{\mu^b R^{bN}}{N^b}$ .

Thus we can write

$$\begin{aligned} I_1 &\geq -A(\epsilon) + \frac{N}{b} (1 - \frac{1}{N}) \int_0^{R^N} s^{-\frac{1}{N}} (s + \epsilon)^{-a} w^b ds \\ &\quad - a \frac{N}{b} \int_0^{R^N} s^{-\frac{1}{N}} (s + \epsilon)^{-a} w^b ds \\ &= -A(\epsilon) - \frac{N}{b} [a - (1 - \frac{1}{N})] \int_0^{R^N} s^{-\frac{1}{N}} (s + \epsilon)^{-a} w^b ds. \end{aligned} \tag{2.13}$$

Here, taking into account of (2.10) and (2.15) we observe that  $\gamma < 1 - \frac{2}{N} < 1 - \frac{1}{N} < a$ , so that  $a - \gamma > 0$ , and hence we can define  $\delta > 0$  such that  $\delta := \frac{N k_f \bar{C}}{2(b+1)}(a - \gamma)$ . Since  $a > 1$  by (2.10), using the Young inequality in the last term of (2.13), we have

$$\begin{aligned} I_1 &\geq -A(\epsilon) - C_\delta \int_0^{R^N} s^{-\frac{b+1}{N}} (s + \epsilon)^{-a(b+1)+b(a+1-\gamma)} ds \\ &\quad - \frac{b}{b+1} \delta \int_0^{R^N} (s + \epsilon)^{-(a+1-\gamma)} w^{b+1} ds \geq -A(\epsilon) \\ &\quad - C_\delta \int_0^{R^N} s^{-\frac{b+1}{N}} (s + \epsilon)^{-a+b-b\gamma} ds - \delta \int_0^{R^N} (s + \epsilon)^{-(a+1-\gamma)} w^{b+1} ds, \end{aligned} \tag{2.14}$$

where

$$\gamma := 2(1 - \frac{1}{N})\alpha < 2(1 - \frac{1}{N}) \frac{N-2}{2(N-1)} = \frac{N-2}{N} < 1, \tag{2.15}$$

and

$$C_\delta := \frac{1}{b+1} \left\{ \frac{N}{b} [a - (1 - \frac{1}{N})] \right\}^{b+1} \delta^{-b}. \tag{2.16}$$

We now estimate  $I_2$ . Integrating by parts, we have

$$\begin{aligned} I_2 &= N k_f \bar{C} \int_0^{R^N} s^{2(1-\frac{1}{N})\alpha} (s + \epsilon)^{-a} w^b w_s ds \\ &= \frac{N k_f \bar{C}}{b+1} [(s^\gamma (s + \epsilon)^{-a} w^{b+1})]_0^{R^N} - \frac{N k_f \bar{C}}{b+1} \int_0^{R^N} \frac{d}{ds} (s^\gamma (s + \epsilon)^{-a}) w^{b+1} ds \\ &\geq -\frac{N k_f \bar{C}}{b+1} \int_0^{R^N} [\gamma s^{\gamma-1} (s + \epsilon)^{-a} - a s^\gamma (s + \epsilon)^{-a-1}] w^{b+1} ds \end{aligned} \tag{2.17}$$

A similar calculation of  $I_3$  of (2.12) leads to

$$\begin{aligned}
 I_3 &= -\mu\chi k_f \int_0^{R^N} s(s+\epsilon)^{-a} w^{b-1} w_s ds = -\frac{\mu\chi k_f}{b} [s(s+\epsilon)^{-a} w^b]_0^{R^N} \\
 &+ \frac{\mu\chi k_f}{b} \int_0^{R^N} \frac{d}{ds} (s(s+\epsilon)^{-a}) w^b ds = -\frac{\mu\chi k_f}{b} R^N (R^N + \epsilon)^{-a} \frac{\mu^b R^{bN}}{N^b} \\
 &+ \frac{\mu\chi k_f}{b} \int_0^{R^N} [(s+\epsilon)^{-a} - as(s+\epsilon)^{-a-1}] w^b ds \\
 &\geq -A_1(\epsilon) - \frac{\mu\chi k_f}{b} (a-1) \int_0^{R^N} (s+\epsilon)^{-a} w^b ds,
 \end{aligned}
 \tag{2.18}$$

with  $A_1(\epsilon) = \frac{\mu^{1+b}\chi k_f}{b} R^N (R^N + \epsilon)^{-a} \frac{R^{bN}}{N^b}$ .

By replacing (2.14), (2.17) and (2.18) into (2.12) we have

$$\begin{aligned}
 &\frac{1}{b} \frac{d}{dt} \int_0^{R^N} (s+\epsilon)^{-a} w^b(s,t) ds \\
 &\geq -C_\delta \int_0^{R^N} s^{-\frac{b+1}{N}} (s+\epsilon)^{-(a+b-b\gamma)} ds - \delta \int_0^{R^N} (s+\epsilon)^{-(a+1-\gamma)} w^{b+1} ds \\
 &- \frac{Nk_f\bar{C}}{b+1} \int_0^{R^N} [\gamma s^{\gamma-1} (s+\epsilon)^{-a} - as^\gamma (s+\epsilon)^{-a-1}] w^{b+1} ds \\
 &- \frac{\mu\chi k_f}{b} (a-1) \int_0^{R^N} (s+\epsilon)^{-a} w^b ds - (A(\epsilon) + A_1(\epsilon)).
 \end{aligned}$$

Integrating from 0 to  $t \in (0, T_{max})$  we arrive to

$$\begin{aligned}
 &\frac{1}{b} \int_0^{R^N} (s+\epsilon)^{-a} w^b(s,t) ds - \frac{1}{b} \int_0^{R^N} (s+\epsilon)^{-a} w_0^b(s) ds \\
 &\geq -C_\delta t \int_0^{R^N} s^{-\frac{b+1}{N}} (s+\epsilon)^{-(a+b-b\gamma)} ds - \delta \int_0^t \int_0^{R^N} (s+\epsilon)^{-(a+1-\gamma)} w^{b+1} ds d\tau \\
 &- \frac{Nk_f\bar{C}}{b+1} \int_0^{R^N} [\gamma s^{\gamma-1} (s+\epsilon)^{-a} - as^\gamma (s+\epsilon)^{-a-1}] w^{b+1} ds \\
 &- \frac{\mu\chi k_f}{b} (a-1) \int_0^t \int_0^{R^N} (s+\epsilon)^{-a} w^b ds d\tau - \tilde{A}_\epsilon t,
 \end{aligned}$$

where  $\tilde{A}_\epsilon = A(\epsilon) + A_1(\epsilon)$ .

Now, from the monotone convergence theorem, taking  $\epsilon \searrow 0$ , we arrive to

$$\begin{aligned}
 &\frac{1}{b} \int_0^{R^N} s^{-a} w^b(s,t) ds - \frac{1}{b} \int_0^{R^N} s^{-a} w_0^b(s) ds \\
 &\geq -C_\delta t \int_0^{R^N} s^{b-b\gamma-\frac{b+1}{N}-a} ds + \left( \frac{Nk_f\bar{C}}{b+1} (a-\gamma) - \delta \right) \int_0^t \int_0^{R^N} s^{\gamma-a-1} w^{b+1} ds d\tau \\
 &- \frac{\mu\chi k_f}{b} (a-1) \int_0^t \int_0^{R^N} s^{-a} w^b ds d\tau - \tilde{A}t,
 \end{aligned}$$

with  $\tilde{A} = \lim_{\epsilon \searrow 0} \tilde{A}_\epsilon$ .

Note that  $\beta := (b-b\gamma - \frac{b+1}{N} - a) + 1 > 0$  by (2.10), so that  $\int_0^{R^N} s^{b-b\gamma-\frac{b+1}{N}-a} ds = \frac{R^{N\beta}}{\beta}$ . Hence we have

$$\begin{aligned}
 &\frac{1}{b} \int_0^{R^N} s^{-a} w^b(s,t) ds - \frac{1}{b} \int_0^{R^N} s^{-a} w_0^b(s) ds \\
 &\geq \left( \frac{Nk_f\bar{C}}{b+1} (a-\gamma) - \delta \right) \int_0^t \int_0^{R^N} s^{\gamma-a-1} w^{b+1} ds d\tau \\
 &- \frac{\mu\chi k_f}{b} (a-1) \int_0^t \int_0^{R^N} s^{-a} w^b ds d\tau - \tilde{A}t
 \end{aligned}
 \tag{2.19}$$

with

$$\tilde{A} := \tilde{A} + C_\delta \frac{R^{N\beta}}{\beta} = R^{N-1-a} \frac{\mu^b}{b} \frac{R^{bN}}{N^b} [N + \mu\chi k_f R] + C_\delta \frac{R^{N\beta}}{\beta}.$$

Here, recalling the definition  $\delta := \frac{Nk_f\bar{C}}{2(b+1)}(a-\gamma)$ , we note that  $\frac{Nk_f\bar{C}}{b+1}(a-\gamma) - \delta > 0$ .

To estimate the term  $\int_0^{R^N} s^{\gamma-a-1} w^{b+1} ds$  we apply the Hölder inequality:

$$\begin{aligned} \int_0^{R^N} s^{-a} w^b ds &= \int_0^{R^N} s^{-a+\frac{b}{b+1}(a+1-\gamma)} \left( s^{\gamma-a-1} w^{b+1} \right)^{\frac{b}{b+1}} ds \\ &\leq \left( \int_0^{R^N} s^{-a(b+1)+b(a+1-\gamma)} ds \right)^{\frac{1}{b+1}} \left( \int_0^{R^N} s^{\gamma-a-1} w^{b+1} ds \right)^{\frac{b}{b+1}} \\ &= \left( \int_0^{R^N} s^{b-b\gamma-a} ds \right)^{\frac{1}{b+1}} \left( \int_0^{R^N} s^{\gamma-a-1} w^{b+1} ds \right)^{\frac{b}{b+1}} \\ &= \bar{c}^{-1} \left( \int_0^{R^N} s^{\gamma-a-1} w^{b+1} ds \right)^{\frac{b}{b+1}} \end{aligned}$$

with  $\bar{c} = \frac{b-b\gamma-a+1}{R^N(b-b\gamma-a+1)}$  and  $b-b\gamma-a > -1$  (since  $b-b\gamma-a > b-b\gamma-\frac{b+1}{N}-a > -1$ ). From which we derive

$$\int_0^{R^N} s^{\gamma-a-1} w^{b+1} ds \geq \bar{c} \left( \int_0^{R^N} s^{-a} w^b ds \right)^{\frac{b+1}{b}}. \tag{2.20}$$

Substituting (2.20) in (2.19) we obtain

$$\begin{aligned} \frac{1}{b} \int_0^{R^N} s^{-a} w^b(s, t) ds &\geq \frac{1}{b} \int_0^{R^N} s^{-a} w_0^b(s) ds - \frac{\mu \chi k_f}{b} (a-1) \int_0^t \int_0^{R^N} s^{-a} w^b ds d\tau \\ &\quad + \left( \frac{N k_f \bar{C}}{b+1} (a-\gamma) - \delta \right) \bar{c} \int_0^t \left( \int_0^{R^N} s^{-a} w^b ds \right)^{\frac{b+1}{b}} d\tau - \bar{A}t, \end{aligned}$$

i.e. (2.11) with  $\delta_1 := \frac{\mu \chi k_f}{b} (a-1)$ ,  $\delta_2 := \left( \frac{N k_f \bar{C}}{b+1} (a-\gamma) - \delta \right) \bar{c}$  and  $\delta_3 := \bar{A}$ .  $\square$

**Proof of Theorem 1.1.** Set  $y(t) = \int_0^{R^N} s^{-a} w^b ds$ . In light of Lemma 2.3, we notice that the inequality (2.11) can be rewritten in the following form :

$$y(t) \geq \delta_0 - \delta_1 \int_0^t y(\tau) d\tau + \delta_2 \int_0^t y^{1+d}(\tau) d\tau - \delta_3 t, \quad \forall t \in (0, T_{max}), \tag{2.21}$$

with  $d = \frac{1}{b}$  and  $\delta_0 = \frac{1}{b} \int_0^{R^N} s^{-a} w_0^b(s) ds$ .

On the other hand, there exists  $\epsilon_1 \in (0, \delta_2/\delta_1)$  such that

$$y(t) = \int_0^{R^N} s^{-a} w^b ds \leq \epsilon_1 \left( \int_0^{R^N} s^{-a} w^b ds \right)^{1+d} + c(\epsilon_1) = \epsilon_1 y^{1+d}(t) + c(\epsilon_1). \tag{2.22}$$

Using (2.22) in (2.21) we obtain

$$y(t) \geq \delta_0 + C_3 \int_0^t y^{1+d}(\tau) d\tau - C_4 t, \quad \forall t \in (0, T_{max}),$$

where  $C_3 := \delta_2 - \delta_1 \epsilon_1 > 0$  (due to the choice of  $\epsilon_1$ ),  $C_4 := \delta_1 c(\epsilon_1) + \delta_3$ .

Let us introduce the function

$$z(t) = \delta_0 - \left( \frac{C_4}{C_3} \right)^{\frac{1}{1+d}} + C_3 \int_0^t y^{1+d} d\tau - C_4 t,$$

which satisfies

$$\begin{aligned} z'(t) &= C_3 y^{1+d}(t) - C_4 \geq C_3 \left( z(t) + \left( \frac{C_4}{C_3} \right)^{\frac{1}{1+d}} \right)^{1+d} - C_4 \geq C_3 z^{1+d}(t), \\ z(0) &= \delta_0 - \left( \frac{C_4}{C_3} \right)^{\frac{1}{1+d}}. \end{aligned} \tag{2.23}$$

Following the step in the proof of Theorem 0.1 in [20] we can conclude that  $z(0) = \delta_0 - \left( \frac{C_4}{C_3} \right)^{\frac{1}{1+d}} > 0$  for the hypothesis on initial data. In fact, let us introduce the following nonnegative function

$$\psi_\epsilon(s) := \frac{\mu}{N} \frac{R^N + \epsilon}{s + \epsilon} s, \quad s \in [0, R^N], \quad \epsilon > 0,$$

satisfying  $\psi_\epsilon \nearrow \frac{\mu R^N}{N}$  as  $\epsilon \searrow 0$ . From the monotone convergence theorem, since  $a > 1$ , we obtain

$$\int_0^{R^N} s^{-a} \psi_\epsilon^b(s) ds \rightarrow \infty, \quad \text{as } \epsilon \searrow 0.$$

Finally, for some sufficiently small  $\epsilon > 0$  we define  $w_0(s) := \psi_\epsilon(s)$ ,  $s \in [0, R^N]$  and we easily obtain  $z(0) > 0$ .



We observe that  $w_0$  belongs to  $C^\infty([0, R^N])$  with  $w_0(0) = 0$ ,  $w_0(R^N) = \frac{\mu R^N}{N}$  and  $w_{0s}(s) > 0$  for all  $s \in [0, R^N]$ . As a consequence  $u_0(x) := Nw_{0s}(|x|^N)$  for  $x \in \bar{\Omega}$  is radially symmetric, smooth and positive in  $\bar{\Omega}$  with  $\frac{1}{|\Omega|} \int_{\Omega} u_0 dx = \mu$ .

Since  $z(0) > 0$ , by comparison theorem  $z(t)$  is positive. The inequality (2.23) is equivalent to

$$-\frac{1}{d} \left( z^{-d}(t) \right)' \geq C_3,$$

from which, integrating from 0 to  $t$ , we obtain

$$-z^{-d}(0) \leq z^{-d}(t) - z^{-d}(0) \leq -dC_3t, \quad t \in (0, T_{max}),$$

and  $t \leq \frac{z^{-d}(0)}{dC_3}$ . If  $t \rightarrow T_{max}$  then  $z \rightarrow +\infty$  with  $T_{max} \leq \frac{z^{-d}(0)}{dC_3}$ .

The proof of Theorem 1.1 is completed by the blow-up criterion in Lemma 2.1.  $\square$

### 3. Blow-up in $L^p$ -norm

In this section we prove by contradiction that  $u(x, t)$  blows up in  $L^p(\Omega)$ -norm for all  $p > N$ . We have the following lemma.

**Lemma 3.1.** *Let  $\Omega \subset \mathbb{R}^N$ ,  $N \geq 3$  be a bounded and smooth domain. Let  $(u, v)$  be a classical solution of system (1.1). Then if there exists  $C > 0$  such that for some  $p_0 > N$ ,*

$$\|u(\cdot, t)\|_{L^{p_0}(\Omega)} \leq C \quad \text{for any } t \in (0, T_{max}),$$

then, for some  $\hat{C} > 0$ ,

$$\|u(\cdot, t)\|_{L^\infty(\Omega)} \leq \hat{C} \quad \text{for any } t \in (0, T_{max}).$$

**Proof.** Let us consider the  $L^p$ -norm of  $u(\cdot, t)$  of (1.1) for all  $p > \max\{p_0, 2N\}$ . Using the first equation of (1.1) and then integrating by parts, we obtain

$$\begin{aligned} \frac{1}{p} \frac{d}{dt} \int_{\Omega} u^p dx &= \int_{\Omega} u^{p-1} u_t dx \\ &= - \int_{\Omega} \nabla u^{p-1} \frac{u \nabla u}{\sqrt{u^2 + |\nabla u|^2}} dx + \chi k_f \int_{\Omega} \nabla u^{p-1} \frac{u \nabla v}{(1 + |\nabla v|^2)^\alpha} dx \\ &= -(p-1) \int_{\Omega} \frac{u^{p-1} |\nabla u|^2}{\sqrt{u^2 + |\nabla u|^2}} dx + \chi k_f (p-1) \int_{\Omega} u^{p-1} \frac{\nabla u \nabla v}{(1 + |\nabla v|^2)^\alpha} dx. \end{aligned}$$

We use the following inequality proved by Bellomo and Winkler in [1, Lemma 6.1]:

$$\int_{\Omega} u^{p-1} |\nabla u| dx \leq \int_{\Omega} \frac{u^{p-1} |\nabla u|^2}{\sqrt{u^2 + |\nabla u|^2}} dx + \int_{\Omega} u^p dx.$$

Thus we have

$$\frac{1}{p} \frac{d}{dt} \int_{\Omega} u^p dx \leq -(p-1) \int_{\Omega} u^{p-1} |\nabla u| dx + (p-1) \int_{\Omega} u^p dx + \mathcal{J}, \tag{3.1}$$

for all  $t \in (0, T_{max})$ , where

$$\mathcal{J} := \chi k_f (p-1) \int_{\Omega} u^{p-1} \frac{\nabla u \nabla v}{(1 + |\nabla v|^2)^\alpha} dx.$$

Integrating by parts and using the boundary conditions and the second equation of (1.1), we obtain

$$\begin{aligned} \mathcal{J} &= \chi k_f \frac{p-1}{p} \int_{\Omega} \frac{\nabla u^p \nabla v}{(1 + |\nabla v|^2)^\alpha} dx \\ &= -\chi k_f \frac{p-1}{p} \int_{\Omega} \frac{u^p \Delta v}{(1 + |\nabla v|^2)^\alpha} dx + \alpha \chi k_f \frac{p-1}{p} \int_{\Omega} u^p \frac{\nabla v \cdot \nabla (|\nabla v|^2)}{(1 + |\nabla v|^2)^{\alpha+1}} dx \\ &= -\chi k_f \frac{p-1}{p} \int_{\Omega} \frac{u^p (\mu - u)}{(1 + |\nabla v|^2)^\alpha} dx + \alpha \chi k_f \frac{p-1}{p} \int_{\Omega} u^p \frac{\nabla v \cdot \nabla (|\nabla v|^2)}{(1 + |\nabla v|^2)^{\alpha+1}} dx \\ &\leq \chi k_f \frac{p-1}{p} \int_{\Omega} u^{p+1} dx + \alpha \chi k_f \frac{p-1}{p} \int_{\Omega} u^p \frac{\nabla v \cdot \nabla (|\nabla v|^2)}{(1 + |\nabla v|^2)^{\alpha+1}} dx, \end{aligned} \tag{3.2}$$

where in the last step we neglected the negative term  $-\mu \chi k_f \frac{p-1}{p} \int_{\Omega} u^p dx$  and we used the inequality  $\frac{1}{(1 + |\nabla v|^2)^\alpha} \leq 1$  as  $\alpha > 0$ .

To estimate the second term in (3.2) we follow the step in the proof of [12, (5.7)] and by using the radially symmetric setting we can obtain

$$\mathcal{J} \leq 2\mu \alpha \chi k_f \frac{p-1}{p} \int_{\Omega} u^p dx + \chi k_f \frac{p-1}{p} \left( 1 + 2\alpha N(N-1) \bar{c}_1 \right) \int_{\Omega} u^{p+1} dx \tag{3.3}$$

$$+ 2\alpha N(N-1)\bar{c}_2 \chi k_f \frac{p-1}{p} \left( \int_{\Omega} u^{p+1+\epsilon} dx \right)^{\frac{p+1}{p+1+\epsilon}}$$

with some constants  $\bar{c}_1, \bar{c}_2 > 0$  depending on small  $\epsilon > 0$ .

Taking into account of (3.3) in (3.1), we can write

$$\begin{aligned} \frac{1}{p} \frac{d}{dt} \int_{\Omega} u^p dx &\leq -\frac{p-1}{p} \int_{\Omega} |\nabla u^p| dx \\ &+ \mathcal{A} \int_{\Omega} u^p dx + \mathcal{B} \int_{\Omega} u^{p+1} dx + C \left( \int_{\Omega} u^{p+1+\epsilon} dx \right)^{\frac{p+1}{p+1+\epsilon}} \end{aligned} \tag{3.4}$$

for all  $t \in (0, T_{max})$ , where

$$\begin{cases} \mathcal{A} &= p-1 + 2\mu\alpha \chi k_f \frac{p-1}{p}, \\ \mathcal{B} &= \chi k_f \frac{p-1}{p} (1 + 2\alpha N(N-1)\bar{c}_1), \\ \mathcal{C} &= 2\alpha N(N-1)\bar{c}_2 \chi k_f \frac{p-1}{p}. \end{cases} \tag{3.5}$$

By using the Gagliardo–Nirenberg inequality and the Young inequality we estimate the term  $\int_{\Omega} u^{p+1} dx$  as

$$\begin{aligned} \int_{\Omega} u^{p+1} dx &= \|u^p\|_{L^1(\Omega)} \leq C_{GN} \|\nabla u^p\|_{L^1(\Omega)}^\theta \|u^p\|_{L^{\frac{1}{2}}(\Omega)}^{1-\theta} + C_{GN} \|u^p\|_{L^{\frac{1}{2}}(\Omega)}, \\ &\leq C_{GN} \left( \int_{\Omega} |\nabla u^p| dx \right)^\theta \left( \int_{\Omega} u^{\frac{p}{2}} dx \right)^{2(1-\theta)} + C_{GN} \left( \int_{\Omega} u^{\frac{p}{2}} dx \right)^2 \\ &\leq \epsilon_1 \int_{\Omega} |\nabla u^p| dx + c(\epsilon_1) \left( \int_{\Omega} u^{\frac{p}{2}} dx \right)^2, \end{aligned} \tag{3.6}$$

where  $\theta = \frac{N}{N+1} \in (0, 1)$  and  $\epsilon_1 > 0$  to be fixed later. Similarly, we obtain

$$\begin{aligned} \int_{\Omega} u^{p+1} dx &= \|u^p\|_{L^{\frac{p+1}{p}}(\Omega)}^{\frac{p+1}{p}} \leq C_{GN} \|\nabla u^p\|_{L^1(\Omega)}^{\frac{p+1}{p}\bar{\theta}} \|u^p\|_{L^{\frac{1}{2}}(\Omega)}^{\frac{p+1}{p}(1-\bar{\theta})} + C_{GN} \|u^p\|_{L^{\frac{1}{2}}(\Omega)}^{\frac{p+1}{p}} \\ &\leq \epsilon_2 \int_{\Omega} |\nabla u^p| dx + c(\epsilon_2) \left( \int_{\Omega} u^{\frac{p}{2}} dx \right)^{\frac{2(p+1-N)}{p-2N}} + C_{GN} \left( \int_{\Omega} u^{\frac{p}{2}} dx \right)^{\frac{2(p+1)}{p}} \end{aligned} \tag{3.7}$$

where  $\bar{\theta} = \frac{1+\frac{1}{p+1}}{1+\frac{1}{N}} = \frac{N(p+2)}{(N+1)(p+1)} \in (0, 1)$  and  $\epsilon_2 > 0$  to be fixed later. Also, we have

$$\begin{aligned} \left( \int_{\Omega} u^{p+1+\epsilon} dx \right)^{\frac{p+1}{p+1+\epsilon}} &= \|u^p\|_{L^{\frac{p+1+\epsilon}{p}}(\Omega)}^{\frac{p+1}{p}} \\ &\leq C_{GN} \|\nabla u^p\|_{L^1(\Omega)}^{\frac{\bar{\theta} p+1}{p}} \|u^p\|_{L^{\frac{1}{2}}(\Omega)}^{(1-\bar{\theta})\frac{p+1}{p}} + C_{GN} \|u^p\|_{L^{\frac{1}{2}}(\Omega)}^{\frac{p+1}{p}} \\ &\leq \epsilon_3 \int_{\Omega} |\nabla u^p| dx + c(\epsilon_3) \left( \int_{\Omega} u^{\frac{p}{2}} dx \right)^\beta + C_{GN} \left( \int_{\Omega} u^{\frac{p}{2}} dx \right)^{\frac{2(p+1)}{p}}, \end{aligned} \tag{3.8}$$

where  $\tilde{\theta} = \frac{1+\frac{1+\epsilon}{p+1+\epsilon}}{1+\frac{1}{N}} = \frac{N(p+2+2\epsilon)}{(N+1)(p+1+\epsilon)} \in (0, 1)$  and  $\tilde{\theta} \frac{p+1}{p} \in (0, 1)$  for sufficiently small  $\epsilon > 0$ , and  $\beta = \frac{2(p+1)(p+1-N)-(N-1)\epsilon}{(p+1)(p-2N)-((N-1)p+2N)\epsilon}$  and  $\epsilon_3 > 0$  to be fixed later.

Substituting (3.6), (3.7), (3.8) in (3.4), and noting that  $2 < \frac{2(p+1)}{p} < \frac{2(p+1-N)}{p-2N} < \beta$ , we infer from the Young inequality that

$$\begin{aligned} \frac{1}{p} \frac{d}{dt} \int_{\Omega} u^p dx &\leq -\left( \frac{p-1}{p} - \mathcal{A}\epsilon_1 - \mathcal{B}\epsilon_2 - \mathcal{C}\epsilon_3 \right) \int_{\Omega} |\nabla u^p| dx \\ &+ c_1(\epsilon_1, \epsilon_2, \epsilon_3) \left( \int_{\Omega} u^{\frac{p}{2}} dx \right)^2 + c_2(\epsilon_1, \epsilon_2, \epsilon_3) \left( \int_{\Omega} u^{\frac{p}{2}} dx \right)^\beta. \end{aligned}$$

Taking  $\epsilon_1, \epsilon_2, \epsilon_3$  suitably small and using (3.6) in the form

$$\int_{\Omega} |\nabla u^p| dx \geq c_4 \int_{\Omega} u^p dx - c_5 \left( \int_{\Omega} u^{\frac{p}{2}} dx \right)^2,$$

we arrive at

$$\frac{d}{dt} \int_{\Omega} u^p dx + \int_{\Omega} u^p dx \leq c_6 p^{1+N} \left( \int_{\Omega} u^{\frac{p}{2}} dx \right)^2 + c_7 p \left( \int_{\Omega} u^{\frac{p}{2}} dx \right)^\beta. \tag{3.9}$$

Now we define

$$p := p_k = p_0 2^k, \tag{3.10}$$

for nonnegative integers  $k$  and let introduce

$$M_k := \sup_{t \in (0, T)} \int_{\Omega} u^{p_k} dx, \tag{3.11}$$

with  $k \geq 1$  and  $T \in (0, T_{max})$  is fixed.

From the definition of  $p$  in (3.10) and of  $M_k$  in (3.11) we have

$$\int_{\Omega} u^{\frac{p}{2}} dx = \int_{\Omega} u^{p_0 2^{k-1}} \leq M_{k-1}, \quad t \in (0, T).$$

This in conjunction with (3.9) implies that

$$\frac{d}{dt} \int_{\Omega} u^{p_k} dx + \int_{\Omega} u^{p_k} dx \leq c_6 p_k^{1+N} M_{k-1}^2 + c_7 p_k M_{k-1}^{\beta}.$$

By comparison arguments we obtain

$$\int_{\Omega} u^{p_k} dx = M_k \leq \max \left\{ \int_{\Omega} u_0^{p_k} dx, c_6 p_k^{1+N} M_{k-1}^2 + c_7 p_k M_{k-1}^{\beta} \right\} \text{ for all } k \geq 1.$$

Now, if there exists a sequence  $(k_j)_{j \in \mathbb{N}} \subset \mathbb{N}$  such that  $k_j \rightarrow \infty$  as  $j \rightarrow \infty$  and

$$M_{k_j} \leq \int_{\Omega} u_0^{p_{k_j}} dx \text{ for all } j \in \mathbb{N},$$

then we have

$$\sup_{t \in (0, T)} \|u(\cdot, t)\|_{L^{p_{k_j}}(\Omega)} \leq \|u_0\|_{L^{p_{k_j}}(\Omega)},$$

and taking  $j \rightarrow \infty$ , we obtain

$$\sup_{t \in (0, T)} \|u(\cdot, t)\|_{L^{\infty}(\Omega)} \leq \|u_0\|_{L^{\infty}(\Omega)}.$$

Conversely, if no such sequence exists, then for some large  $k_0 \in \mathbb{N}$  we have

$$M_k \leq c_6 p_k^{1+N} M_{k-1}^2 + c_7 p_k M_{k-1}^{\beta} \text{ for all } k \geq k_0.$$

Since  $p_k = p_0 2^k$  and  $2 < \beta \leq 2 + \frac{c_8}{2^k}$  for all  $k \geq k_0$  with some large  $k_0 \geq 1$ , there exists a number  $\delta > 1$  independent of  $T$  such that

$$M_k \leq \delta^k M_{k-1}^{2 + \frac{c_8}{2^k}} \text{ for all } k \geq k_0,$$

and by induction and limiting procedure as in [21, p. 714] we obtain

$$\limsup_{k \rightarrow \infty} M_k^{\frac{1}{p_k}} \leq c_9$$

for some  $c_9 > 0$ . In view of the definition (3.11), this proves

$$\sup_{t \in (0, T)} \|u(\cdot, t)\|_{L^{\infty}(\Omega)} \leq c_9.$$

Thus in both cases we obtain that if  $\|u(\cdot, t)\|_{L^{p_0}(\Omega)}$  is bounded for some  $p_0 > N$  then  $\|u(\cdot, t)\|_{L^{\infty}(\Omega)}$  is bounded.  $\square$

**Proof of Theorem 1.2.** Taking into account of Lemma 3.1, if  $\|u(\cdot, t)\|_{L^{p_0}(\Omega)} \leq C$  for some  $p_0 > N$ , then also  $\|u(\cdot, t)\|_{L^{\infty}(\Omega)}$  is bounded. This is a contradiction since Theorem 1.1 holds. Therefore  $\limsup_{t \nearrow T_{max}} \|u(\cdot, t)\|_{L^p(\Omega)} = \infty$  for all  $p > N$ .  $\square$

**Remark 3.1.** The investigation on blow-up solutions of system (1.1) goes on with the study of the behavior near the blow-up time  $T_{max}$ . Since it is not always possible to compute  $T_{max}$  we think that deriving a lower bound is a matter of great importance as in [12, Theorem 1.3].

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